# Cycle 25, Space Weather and Propagation for 2023

Carl Luetzelschwab K9LA

e-mail: k9la@arrl.net

website: <a href="https://k9la.us">https://k9la.us</a>



spiral aurora over Tromso, Norway December 11, 2022

#### **Previous Presentations**

- May 2013
  - Cycle 24, disturbances to propagation, several DXpeditions
- February 2017
  - Cycle 24, Cycle 25, prediction software
- March 2023
  - Cycle 25, space weather, propagation for 2023, 10m long path

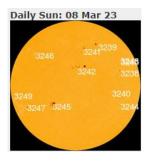
## Agenda

- Cycle 25
  - How is it doing?
- Space weather and propagation
  - Do all those parameters tell us what the ionosphere is doing right now?
- Propagation forecast for 2023
  - What to expect on the bands
- 10-Meter long path
  - Add some spice to your operating this spring thru fall
  - Should apply to 15m and 12m, too
- Bonus slides from a question at the meeting



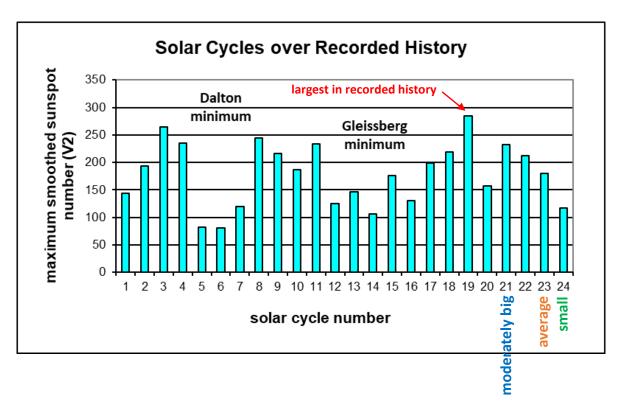
## Cycle 25





## Historical Look at All 24 Cycles

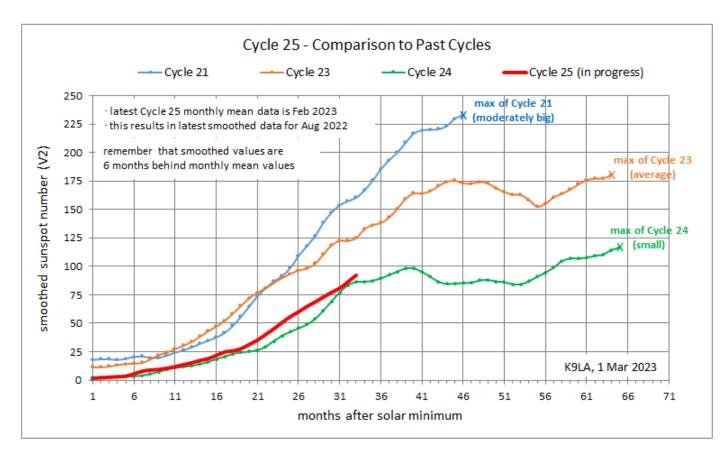
- Cycle 1 began in 1755
  - Maunder Minimum occurred from 1645-1715 with few sunspots
- We've gone through 3 periods of 'big' solar cycles
  - Cycles 1-4, 8-11, 17-23
- We've gone through 2 periods of 'small' solar cycles
  - Cycles 5-7, 12-16
- With Cycle 24, we appear to have entered a third period of small solar cycles



Will Cycle 25 get us out of this possible third period of small cycles?

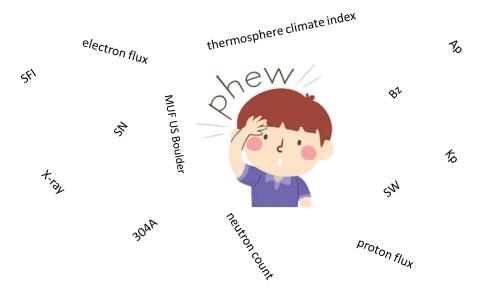
## How Is Cycle 25 Doing?

- Solar minimum was in December 2019
- We currently have 33 months of smoothed sunspot number data since solar minimum
- So far, we appear to be tracking the small Cycle 24
- Thankfully we have enough EUV (extreme ultra-violet) radiation for 15m, 12m and 10m
  - EUV is the true ionizing radiation for the F2 region of the ionosphere
  - Sunspots and 10.7 cm solar flux are <u>proxies</u> for EUV



Hopefully Cycle 25 will move up to an 'average' cycle

## Space Weather and Propagation



## Development of the Model of the Ionosphere

- With solar data and ionosphere data, the result we desired was a correlation between a daily solar parameter (sunspots or 10.7 cm solar flux) and what the ionosphere is doing today
- That didn't happen the correlation was poor
- Why? Because there are <u>three sources of variability</u> of the ionosphere and we only understand two of them
  - Solar radiation sunspots, 10.7 cm solar flux, EUV and MUF US Boulder
  - Geomagnetic field activity K index, A index, Bz and solar wind
  - Events in the lower atmosphere coupling up to the ionosphere no parameters (yet)
- It's tough to make a daily model with data from only 2 of the 3 sources

## Here's What We Ended Up With

- A monthly <u>median</u> model of the ionosphere
  - Median implies 50% probability
  - Correlated to a smoothed solar parameter (long-term average of daily data)
  - Our understanding of the ionosphere is statistical over a month's time frame
- Our propagation predictions give monthly <u>median</u> MUF and signal strength
  - There's a distribution about these median values
  - The MUF on any given day in the month could be somewhat higher to much lower than the median
  - The signal strength on any given day in the month could be somewhat higher to much lower than the median
  - Plugging in the daily 10.7 cm solar flux and the current K index won't make the predictions more accurate

## Parameters That I Consider Important

- Earlier I mentioned SN, SFI, EUV, K,
   A, Bz and solar wind
- One place to get them is from the NØNBH banner at <u>www.qrz.com</u>
- Let's review SN, SFI, EUV, MUF US Boulder, K, A, Bz and solar wind
- Note 'MUF US Boulder' at the bottom

These parameters should give you a general idea of what propagation is like

```
Solar-Terrestrial Data
Provided by N0NBH
          Current Solar Image
```

## Descriptions of Those Parameters

- SFI 10.7 cm solar flux from 65 to 350
- SN sunspot number from 0 to 450
- EUV extreme ultra-violet radiation
- MUF US Boulder MUF when Boulder is midpoint of 3000 km path
- K 3-hr index of the activity of the Earth's magnetic field from 0 to 9 (logarithmic)
- A avg of the eight daily K indices from 0 to 400 (linear)
- Bz north/south component of the IMF\* from +50 to -100
- SW solar wind speed average for quiet time is 400 km/s from 350 km/s to 2000 km/s

We'll focus on the parameters in green

\* IMF is Interplanetary Magnetic Field (Sun's magnetic field)

general correlation

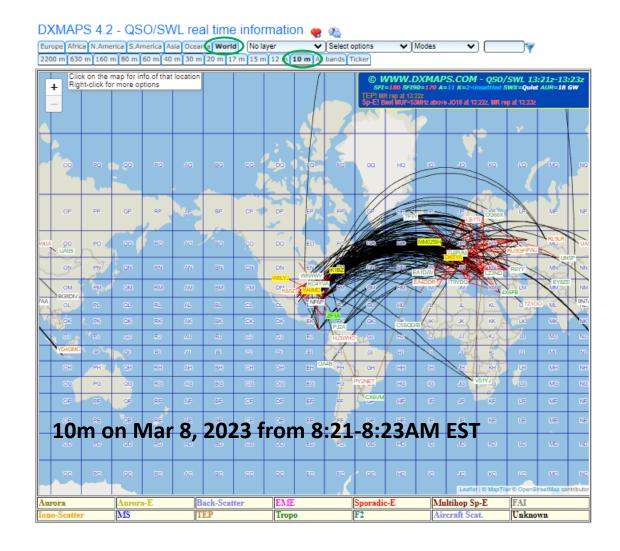
general correlation

### What We Desire for SFI and K

- We need two conditions for a QSO to occur
  - Enough ionization (MUF) to refract the signal back to Earth
  - Low enough loss (absorption, FSPL, antenna gains, transmitter power, receiver MDS, ground reflection loss, local noise) to make signal readable (or detectable)
- What we desire
  - Generally K ≤ 3
    - Exception VHF types like high K indices for propagation via aurora
  - SFI for the higher HF bands (15m, 12m, 10m)
    - 15m: need smoothed SFI > 90
    - 10m: need smoothed SFI > 100
  - Even if SFI is at its minimum, 17m and lower frequencies are still open
- Where we are right now
  - Smoothed SFI ~120

## What Are the Bands Doing Right Now?

- If you don't want to mess with propagation predictions or with all those space weather parameters, go to dxmaps.com
- Select a view (World, NA, . . . )
- Select a band
- Other methods
  - KC2G MUF map (next slide)
  - PSKreporter
  - WSPRnet
  - Reverse Beacon Network
  - IARU/NCDXF beacons

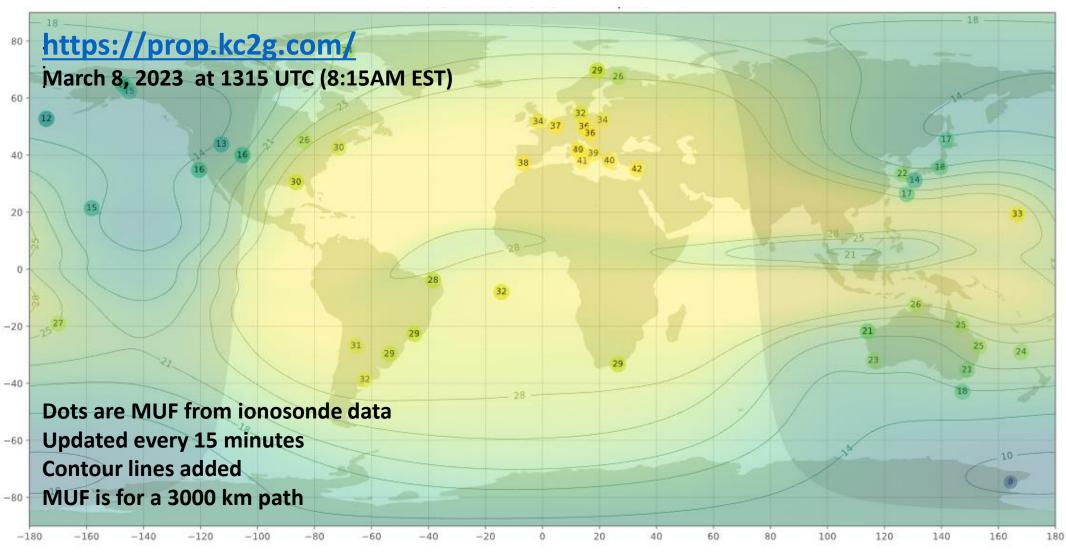


#### Real-Time Assessment of the Bands

for websites on previous slide

- dxmaps.com
- KC2G MUF map <a href="https://prop.kc2g.com/">https://prop.kc2g.com/</a>
- PSKreporter <a href="https://pskreporter.info/pskmap.html">https://pskreporter.info/pskmap.html</a>
- WSPRnet <a href="https://www.wsprnet.org/drupal/wsprnet/map">https://www.wsprnet.org/drupal/wsprnet/map</a>
- Reverse Beacon Network <a href="https://www.reversebeacon.net/">https://www.reversebeacon.net/</a>
- IARU/NCDXF beacons <a href="https://www.ncdxf.org/beacon/">https://www.ncdxf.org/beacon/</a>

## What's the MUF Doing Right Now?



## Propagation Forecast for 2023



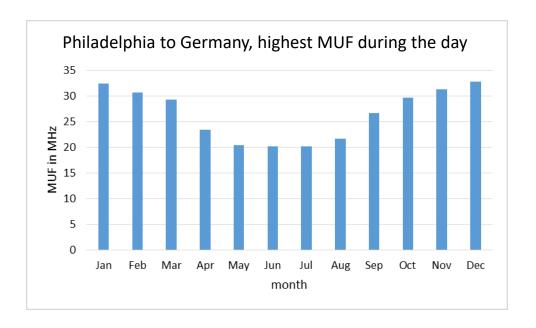
## By the Bands

- 160m, 75m/80m, 60m, 40m (ionospheric absorption is critical)
  - The low bands should be good at night
  - Caveat don't know what to say about 160m should be better than it is

- 30m, 20m, 17m
  - Should be great all year

## By the Bands

- 15m, 12m, 10m (MUF is critical)
  - Should be great for spring contests
  - Should be great for fall/winter contests
    - CQ WW DX in Oct and Nov, et al
  - IARU contest in July will suffer from degraded summer propagation (due to change in atmospheric composition)
    - But watch for sporadic E

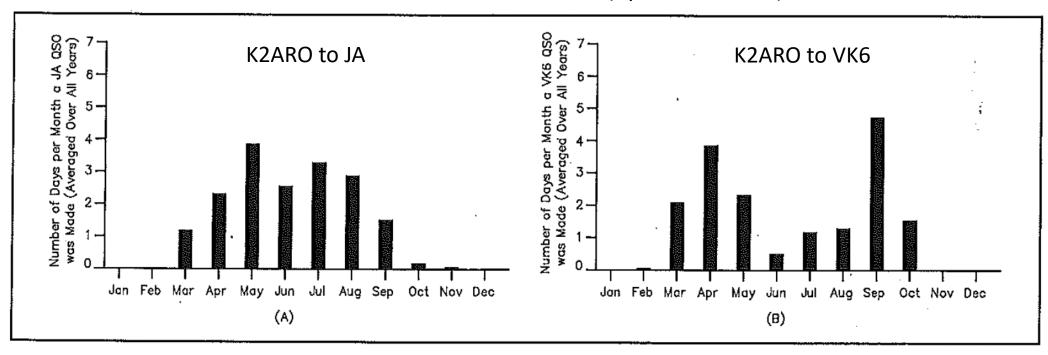


## 10-Meter Long Path

turn your antenna the other way around

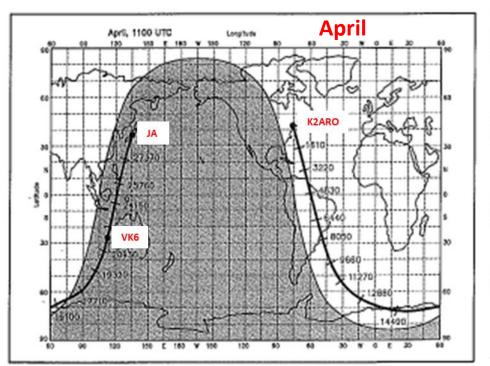
### **K2ARO** Data

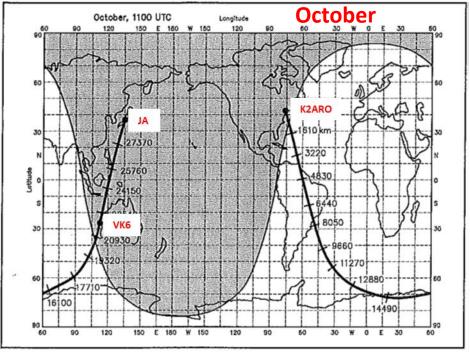
Data from Jan 1979 thru Dec 1993 (Cycles 21 and 22)



- 10m long path available from March thru October
- More pronounced drop out to VK6 in the summer months
  - Suspect that VK6 is more multi-hop whereas JA can have TEP hop

## Typical Long Path

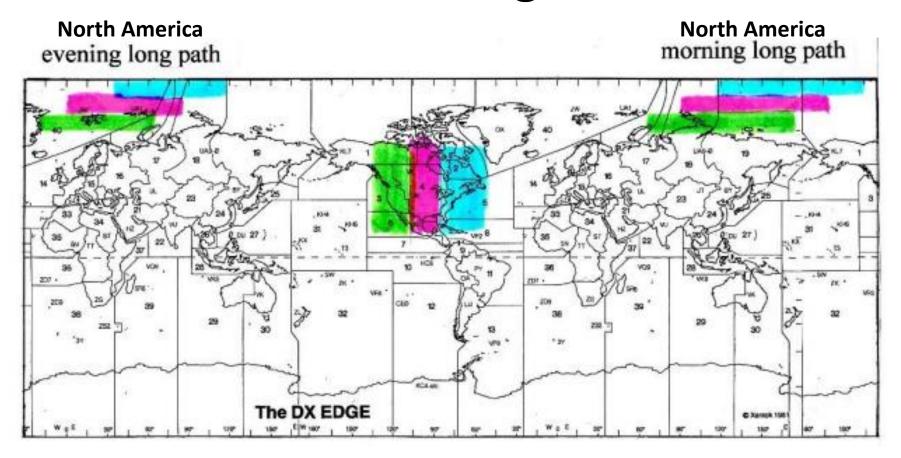




- 10m long path not necessarily a grayline path\*
- Mar 22<sup>nd</sup> to Sep 22<sup>nd</sup> most productive
  - 90% of JA QSOs, 77% of VK6 QSOs
- Before Mar 21<sup>st</sup> and after Sep 23<sup>rd</sup>
  - 10% of JA QSOs, 23% of VK6 QSOs

\* A grayline path is when the entire path aligns with or is very near the terminator

## The Big Picture



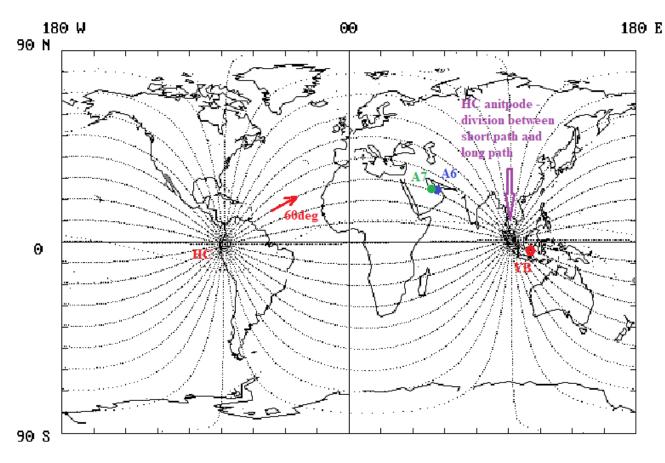
The real question - is anyone on the other end?

For more details: <a href="https://k9la.us/A Refresher on 10m Long Path.pdf">https://k9la.us/A Refresher on 10m Long Path.pdf</a>

Data from logs from K2ARO, N6AV (SK), NT5C (SK) and skeds with JH3DPB (SK)

## HC to A7, A6, YB on 60° Heading on 6m

#### Great-Circle Paths from HC



to	short path	long path
A7	59° / 13988 km	
A6	59° / 14284 km	
YB		64° / 21288 km

- On a heading of 60°
  - A7 is short path
  - A6 is short path
  - YB is long path
- Path stays at low latitudes where the ionosphere is most robust

## Summary

- Cycle 25 is awake and is in its ascent maybe up to an 'average' cycle
- Solar maximum around 2024/2025
- So far it kind of looks like another small cycle
  - We'll either confirm or refute that around 2025
- Even if it is a small cycle, now and around solar maximum will offer worldwide propagation with modest power (100W) and simple antennas (vertical or dipole) on 15m, 12m and 10m
- Should have more 6m F2 propagation this fall/winter
- The digital modes offer an advantage over CW and SSB
  - Can decode a signal farther down in the noise
  - This is a big deal on 10m and 6m where the MUF is critical
- There are tools on the internet to determine what the bands are doing right now